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Variable Wind Ripple Migration at Great Sand Dunes National Park and Preserve, Observed by Timelapse Imaging

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19 Abstract

20 Granule ripples at Great Sand Dunes National Park and Preserve (GSDNPP) were observed using
21 inexpensive digital timelapse cameras over a 70-day period in winter 2010-2011. The ripples migrated
22 during a handful of discrete events - visible ripple movement occurred on only 11 days during the
23 observation period. The movement conditions are documented with hourly and 15-minute records
24 from two nearby weather stations, and by a cup anemometer at the site itself. During the most
25 prominent movement episode, when local winds averaged $\sim 10\text{m/s}$, ripples of several sizes were
26 observed simultaneously and a reciprocal relationship of ripple size and propagation speed was seen,
27 with small ($\sim 10\text{cm}$) ripples moving at $\sim 1.4\text{ cm/min}$, and larger ($\sim 80\text{cm}$) ripples at $\sim 0.15\text{ cm/min}$. Ripple
28 sizes and morphologies evolve throughout the observation period.

29

30

31 Keywords

32 Aeolian ripples; bedform dynamics ; timelapse camera

33

34 1. Introduction

35 The movement and evolution of aeolian bedforms, while dramatically fast by geological standards, occur
36 at rates that are inconveniently slow to observe directly. Furthermore, changes are often strongly
37 episodic, driven by environmental conditions such as windspeed, that only occasionally exceed an action
38 threshold. Thus aeolian process rate measurements demand exceptional patience, and/or luck, from
39 the observer. The approach to measure wind ripple motion has typically (Yitzaq et al., 2008; Zimbelman
40 et al., 2009) been to install markers and to measure motion between visits to a site. Such observations
41 have been important milestones in our understanding of aeolian processes, but suffer from the fact they
42 record the 'integral' movement over what may be a long period during which motion only occurred for a
43 small fraction of the time. Additionally, particularly in regular ripple patterns, observations can be
44 'aliased', wherein one cannot be certain that the same feature is being measured.

45 As discussed in Lorenz (2010, 2011) technological developments mean digital cameras are now available
46 with low enough power consumption and large enough memory capacity to permit automated battery-
47 powered timelapse observation sequences of thousands of images to more comprehensively observe
48 motion and evolution, over anything from hours to months. In particular, in contrast with 'supervised '
49 field experiments of short duration (e.g. Andreotti et al., 2006) such systems are now sufficiently
50 inexpensive as to be 'expendable' - worth risking in extended unattended observations in the field
51 where damage or theft might occur. In this paper, we describe such a field observation of a set of
52 granule ripples at Great Sand Dunes National Park and Preserve (GSDNPP), over a several-week period.
53 Granule ripples in this area have been studied as analogs of ripple features seen on Mars (Zimbelman et
54 al., 2009).

55

56 2. Material and Methods

57 2.1 Field Site

58 The study site (Fig. 1) is a large parabolic dune located at 37°41'35.30"N 105°35'11.25"W. This area is
59 within GSDNPP but is far from the main dunefield and is not in an area frequented by visitors. The site is
60 where the dirt road west from the Lodge on the Desert is blocked by the dune, about 3km from the

61 paved road to the Park. The stoss slopes (on the interior of the U-shaped dune) have many prominent
62 granule ripples.

63 A barbed wire fence (Fig. 2) that runs parallel to the road was buried by the arms of the dune, but is
64 exposed in the center. Three large wooden fenceposts provided convenient mounting points for our
65 instrumentation.

66 < FIGURE 1 - Sketchmap >

67 <FIGURE 2 - site photo >

68

69 The ripples generally have crests (Fig. 3) of granules ~2 mm in diameter. Occasionally, if the ground was
70 damp or frozen, the granules tended to collect entirely as discrete structures with granule-free
71 'interdunes': more generally finer granules were found between crests. Beneath what was typically a
72 monolayer of granules, finer sand (~200 micron) was found.

73 < FIGURE 3 - closeup >

74

75 Granule ripples commonly form at GSDNPP due to the bimodal sand size distribution present there.
76 Most of the sand is fine grained with a medium sand mode. There is a coarse fraction of very coarse
77 sand, granules and pebbles. At the study site, the coarse fraction is found in the trough between the
78 arms of the parabolic dune as wind deflation lowers the area and results in a lag surface. During strong
79 wind events, the very coarse sand and granules are transported on to the nose of the parabolic dune
80 and organize into granule ripples.

81 The sand is sourced in the Sangre de Cristo Mountains, which are adjacent to GSDNPP on the east, and
82 also the more distal San Juan Mountains, which begin 65 km west of Great Sand Dunes. Migration of the
83 dunes is described in Marin et al. (2005). Mineralogical surveys of the dune sand lead to conclusions that
84 the majority of the sand is sourced by volcanic rocks of the San Juan Mountains (Hutchinson, 1968;
85 Wiegand, 1977). The age distribution of zircon in the sand suggests that 70% of the dunefield sand
86 originates in the San Juan Mountains and 30% in the Sangre de Cristos (Madole, 2008). The granules
87 and very coarse sand fraction have a granitic composition. The pebbles are coarse-grained metamorphic

88 and ingenious intrusive rocks. The composition and texture of the coarse sediment fraction suggest
89 that they are sourced by the nearby Sangre de Cristo Mountains, and that streams have flowed in the
90 area. An important agent of local sand transport is Mendano Creek which defines the edge of the main
91 sand dune area.

92

93 2.2 Equipment and Setup

94 We report here on results from two different cameras. First, a modified Brinno Gardenwatchcam™
95 (www.brinno.com) was installed on the post (Fig. 4) on the eastern arm of the dune, amid the ripples
96 under study. This unit is marketed for horticultural timelapse observations and records images acquired
97 with a 1.3 megapixel CMOS imager to an AVI file on a USB memory stick (the 2 GB stick supplied can
98 hold over 15,000 frames). The unit is nominally powered by 4 alkaline AA cells (which typically have a
99 capacity of ~2000 mA-hr), although these do not last long enough to fill the memory stick: we modified
100 the unit to be powered from a separate battery box, in this case using alkaline D-cells (~20,000 mA-hr).

101 < FIGURE 4 - fencepost with camera >

102 The camera ('1') was set at ~2 m height, looking roughly northwards, at an angle of ~30 degrees below
103 horizontal - the scene is thus about 4 m away. Based on prior experience with this camera under winter
104 conditions, we set the image acquisition interval to be 10 minutes, which was expected to allow
105 operation for several months. The unit does not record images in darkness, thus after ~60 images at 10
106 minute spacing in one day, there is a jump to the start of the next day (which is non-deterministic, since
107 it depends on light levels and, thus, on cloud and ground cover). However, the unit records a timestamp
108 on each image.

109 A second camera (hereafter 'Camera 2') was installed about 30 m away on the westernmost fencepost,
110 observing the ripple field in a more or less east-north-eastwards direction (and showing the post with
111 the first camera). This unit was a Wingscapes Plantcam (<http://www.wingscapes.com/>) which writes
112 individual images (high - 2560x1920, medium - 2048x1536, or low-resolution 640x480 pixels) to a 2 GB
113 SD memory card at selectable intervals. This unit, which we used unmodified, has a waterproof casing
114 (like the Gardenwatchcam) which accommodates 4 AA cells. Here we used Lithium AA cells, since these

115 have a higher capacity and tolerate low temperatures better than alkaline battery chemistry. The
116 camera was set to acquire images at 15 minute intervals.

117 We documented wind conditions at the site with an Inspeed (www.inspeed.com) cup anemometer,
118 installed at the top of the first fencepost, also at about 2 m height. This unit indicates winds with a reed
119 switch closure frequency of 2.5 mph per Hz : we counted these pulses over 5 s intervals using a Picaxe
120 18X datalogger (www.rev-ed.co.uk). The average and peak wind speed, and the casing temperature, was
121 recorded at hourly intervals into the unit's on-board memory (although a coding error, discovered after
122 installation, corrupted the average wind measurement; fortunately, the peak gust measurement was
123 not affected). Finally, a Measurement Solutions USB-502 temperature and humidity logger (a unit
124 about the size of a marker pen) was installed next to each camera. A separate cup anemometer, using
125 an electric motor as a transducer, was installed about 15 cm above the ground with a Pace Scientific XR-
126 440M datalogger.

127 During the observation interval, the wire fence was removed. When we retrieved our data on 12
128 January 2011, we observed the western fenceposts to be loose, having been undermined by sand
129 movement: we re-seated one and removed the other.

130

131 3. Results

132 3.1 Equipment Performance

133 Camera 1 appears to have operated satisfactorily, recording images up until the night before it was
134 recovered. The video files were examined in a playback application supplied with the camera (which
135 allows stepping through frame-by-frame). An example excerpt from the movie file showing ripple
136 movement on 21-22 November is shown in the supplemental online material. In addition to inclusion in
137 the online version of this paper, the video files are available at <http://www.lpl.arizona.edu/~rlorenz>

138 < VIDEO >

139

140 Some example images showing evolution of the scene are presented in Fig. 5. Curiously, the video
141 record stops late on 11 January, without images from the morning of recovery. It is possible that the

142 video file was incompletely written, truncating the record (although stopping and starting the camera
143 several times upon recovery, which usually prevents such problems, was performed). A more plausible
144 scenario is that the camera happened to cease operation that night, which saw very cold temperatures
145 (-20°C). Although this was not the first occasion such temperatures were encountered, the probability
146 that the camera will stop when chilled increases with time, as the battery capacity is progressively
147 depleted and thus the voltage droop caused by low temperatures will more likely drop below some
148 threshold.

149 < FIGURE 5 - Montage of Images >

150 Camera 2 operated as intended, until the night of November 21, at which point it ceased operation. This
151 was the first night on which temperatures of -10°C were encountered during the observation period.
152 Battery voltage droop seems not to be implicated, since the relatively fresh lithium batteries should
153 have been in good condition, even at these temperatures. We speculate that the electronics in this unit -
154 with which we have less experience than the Brinno camera - are less tolerant of low temperatures than
155 are those of camera 1.

156 One USB temperature logger was found to have stopped during installation, possibly as a result of a
157 reset caused by its battery being knocked out of place. The other USB logger was successfully
158 interrogated, yielding a good temperature and humidity record.

159 The PICAXE anemometer data logger was interrogated successfully, yielding a case temperature and
160 wind gust history. The XR-440M data appears to have been corrupted; efforts are underway in
161 cooperation with the manufacturer to retrieve these data.

162

163 3.2 Ripple Movement and Evolution from Image Data

164 Visual inspection of the camera 1 record shows ripple migration on only a few days. Using the 80 cm
165 spacing of the marker stakes and the width of the fencepost shadow (the post is 15 cm across, but we
166 take into account the distance-dependent broadening (by a few cm) of the shadow by the 0.5 degree
167 diameter of the sun), the migration distance of ripple features - usually the crest - between image
168 frames can be measured with image analysis tools such as ImageJ. Fig. 8 shows the estimated total
169 movement per day over the observation period.

170 < FIGURE 8 - Ripple Movement over 70 days >

171

172 Over the observation period, the morphology, orientation and size of the ripple pattern varies (see Fig.
173 5). It was seen that the migration rate varied across the scene, either due to some slope effects on
174 saltation or reptation trajectories directly, variation of the wind field, or sediment supply, or all these
175 factors. We will return to these effects shortly in considering data from Camera 2. Small ripples - often
176 superposed on a larger-scale megaripple pattern - were observed to move more quickly than large ones,
177 as in our previous timelapse experiments (Lorenz, 2011).

178 Other features of note include the burial of the ripples (and indeed the ground-level cup anemometer)
179 by snow over the period 16-19 December . A tumbleweed is also seen to transit the area, and animal
180 prints indicate that the equipment attracted the interest - fortunately without damage to the
181 instrumentation - of a fox or coyote on several occasions. During a ripple migration episode on
182 November 28, several of the marker stakes fell down, presumably as a result of aeolian undermining.

183 The Camera 2 record is much shorter and of lower spatial resolution than that of Camera 1. However,
184 the image sequence does capture the ripple movement on 21 November. The more distant view (Fig. 9)
185 sees a larger area of the ripples, and with fewer complicating factors such as foreshortening or the post
186 shadow. We can, therefore, use a semi-automated analysis procedure to efficiently extract migration
187 rates.

188 < FIGURE 9 - view from Camera 2 >

189 First, the images (we use images 578 to 608 of the sequence, spanning 10.03 to 17.33hrs local time -
190 although movement is seen in earlier images that day, the images suffer from glare of the rising sun) are
191 contrast-enhanced to make it easier to track features. In our procedure we read each image into the
192 Interactive Data Language (IDL) and subtract a 20-pixel boxcar averaged image from the original (see Fig.
193 10) - essentially the same as the 'unsharp mask' operator in interactive image analysis tools.

194 < FIGURE 10- stretched image >

195 Then, a line from each enhanced image in the sequence is extracted and inserted into a 'waterfall chart'.
196 Extracted lines from successive images are inserted into progressively lower positions in the chart - thus

197 the horizontal dimension of the chart represents the horizontal dimension of the original scene, while
198 the vertical dimension of the chart corresponds to time. Fixed features in the scene generate vertical
199 lines in the waterfall chart, whereas moving features appear as diagonals, with the slope of the
200 diagonals corresponding to the feature propagation speed across the scene. This 'waterfall chart'
201 procedure was introduced in our earlier work (Lorenz, 2011), although the approach here is improved by
202 the better image quality and by the contrast enhancement. We generate 3 separate waterfall charts for
203 the scene (Fig. 11) , corresponding to different vertical positions in the image, which in turn correspond
204 to different heights on the dune, dominated by different ripple sizes. It is seen that the upper line of
205 the image is dominated by a small-wavelength pattern, which propagates quickly (indeed to the point
206 where the migration during one camera interval approaches the pattern wavelength, leading to aliasing)
207 whereas progressively lower in the image the pattern wavelength increases and the propagation speed
208 declines.

209 < FIGURE 11- Waterfall Chart >

210

211 3.3 Meteorological Measurements

212 While the video record is perhaps of interest in its own right, it is of course more useful if the ripple
213 response to the wind can be related to the wind history itself. In their studies of aeolian ripples at
214 GSDNPP, Zimbleman et al. (2009) use meteorological data from Alamosa airport, 45km away. In fact,
215 more proximate observations exist. Here we use data from two nearby stations. First is the RAWS
216 (Remote Automatic Weather Station) at GSDNPP, operated by the National Park Service, data from
217 which was retrieved from the Western Regional Climate Center
218 (<http://www.raws.dri.edu/wraws/scaF.html>) . The Great Sand Dunes station is at 37° 43' 36" N, 105°
219 30' 39" W, about 10 km to the northeast of the dune site. Daily summary statistics (average wind, peak
220 gust and many other meteorological variables) are available for download as above (data used here was
221 downloaded 1/12/2011). Hourly data can also be obtained.

222 The second station is the Indian Springs Meteorological Station, operated by the US Geological Survey
223 Colorado Water Science Center. This site is at 37°45'50.8" N, 105°37'36.4" W (about 9 km to the
224 northwest of the dune site) and reports only temperature, windspeed and direction, although these

225 data are reported 4 times per hour and are available at <http://waterdata.usgs.gov/co/nwis/> (data used
226 in this paper were downloaded 1/12/2011)

227 It may be noted that the Great Sand Dunes RAWS site (2537 m) is topographically more sheltered than
228 Alamosa airport (2297 m) or the Indian Springs site (2344 m), thus it may be expected to have local
229 slope winds, but may see rather reduced winds driven by regional weather compared with the other
230 sites.

231 The winds recorded at the site during the movement episode on November 21 seem in broad accord
232 (Fig. 12) with those measured at Indian Springs, and with the average and gust data recorded at the
233 RAWS station. Interestingly, although the RAWS gust that day was larger than that we measured at the
234 ripple site, in general our anemometry record indicates higher wind gusts - by some 50% - at the ripple
235 site than those at the RAWS station (Fig. 13). Note that the standard RAWS measurement height is 6.1
236 m, while that for National Weather Service ASOS (Automatic Surface Observing System) sites such as
237 Alamosa airport is nominally 10 m. The Indian Springs measurement height is 2.4 m (W. Walker and R.
238 Crowfoot, USGS, Personal Communication, March 2011). The anemometer on the fencepost at our site
239 was at ~2 m.

240 < FIGURE 12 - Nov 21 Winds >

241 < FIGURE 13 - Correlation Plot >

242

243 4. Discussion

244 4.1 Ripple Migration Rate

245 We determine migration rates from the waterfall chart (figure 11) by converting the horizontal
246 dimension in pixels to physical length, deriving the scale factor from the camera field of view (52°) and
247 distance (31 m); a check against the diameter of the fencepost and other fiducials confirms the scale at
248 ~2.2 cm/pixel. Using ImageJ we pick out a number of features A-O (Fig. 11) and measure their width and
249 spacing, and their propagation speed from the angle across the plot.

250 Results are listed in table 1 and shown in Fig. 14. It is seen that the smallest ripples move most quickly
251 (as would be expected). It is not known with certainty, but seems likely from the imagery and our site

252 inspection that the smaller ripples have a typically smaller particle size. The propagation speed of ~ 1.4
253 cm/min approaches the limit at which we could measure speeds without ambiguity due to aliasing (i.e.
254 during the 15-minute interval between images, a ripple moves one wavelength). On the other hand,
255 larger ripples have propagation with uncertainties that formally allow (permitting the slope on the
256 waterfall plot to be bounded by the extent of the ripple) speeds of zero, although with likely values of
257 ~ 0.15 cm/min. The relationship between ripple size (defined by the width of the bright band in the
258 thresholded image, although we invert the colors in the waterfall chart to improve legibility on the
259 printed page) and propagation speed is a very nearly perfect reciprocal (a best fit in Excel yields an
260 exponent of -1.11 with a correlation coefficient of 0.85). Note that the brightness threshold, chosen to
261 force roughly half of the image pixels to become black, yields a somewhat arbitrary definition of the
262 width of a ripple. Also, for a given unchanging ripple, the perceived threshold width may change as
263 illumination changes over the course of a day. However, this approach is conveniently simple and
264 appears to adequately define ripples that move on timescales much shorter than a day.

265 Correlations of speed against ripple spacing, or width plus spacing (since the structures are not strictly
266 periodic, this width plus spacing metric is not formally a wavelength, but may be more comparable to
267 the scales used in other studies) are poorer, see e.g. Fig. 15.

268 < FIGURE 14 - migration rates vs width >

269 < FIGURE 15 - migration rates vs wavelength >

270 Zimbelman et al. (2009) observed motion of a 10 cm-high granule ripple crest of 10.5 cm over 1380
271 minutes (0.007 cm/min, with winds of ~ 9 m/s), at a site in the main part of the Great Sand Dunes
272 National Park and Preserve, about 5 km away from our site. A 3 cm-high ripple was also measured to
273 move 2.1cm in 109 minutes (0.019 cm/min). We estimate the horizontal size of these ripples from
274 Zimbelman et al. (2009) to be ~ 50 cm and ~ 80 cm, respectively. The overnight measurement of 0.007
275 cm/min probably underestimates the instantaneous propagation speed, as the windspeed dropped
276 below the probable movement threshold for at least part ($\sim 1/4$) of this period.

277 Jerolmack et al. (2006) measure the migration speed of ripples at White Sands National Monument :
278 these ripples were predominantly of gypsum composition. Extrapolating their windspeed vs height
279 measurements (their Fig. 4) to our ~ 2 m measurement height gives a freestream windspeed of ~ 10 m/s.
280 The movement of seven ripples (wavelength ~ 1 m, height ~ 1 cm : the width of the ripple may be

281 estimated from their Fig. 9b to be ~ 20 cm) was measured over 72 minutes, to yield rates of 0.02-0.08
282 m/hr, or 0.03-0.13 cm/min.

283 Ripples in fine Sahara sand on the crest of a Egyptian seif dune (Lorenz, 2011) were observed to move in
284 winds of ~ 10 m/s (although the friction velocity may have been quite high due to the streamline
285 compression at the crest of the dune). The wavelength was ~ 10 cm, and the propagation speed
286 (measured by shorter-range timelapse imagery with a higher cadence than in the present work) was
287 determined to be ~ 3 cm/min. This result is rather in accord with measurements made nearly 50 years
288 ago by Sharp (1963) who recorded migration rates of ripples at the Kelso Dunes in the Mojave desert,
289 California of 0.9 to 8 cm/min in windspeeds of 7 to 18 m/s and identifies a linear correlation between
290 windspeed and migration rate. Seppala and Linde (1978) report migration rates in wind tunnel
291 experiments of 0.15 to 0.6 cm/min with windspeed of 4.3 m/s in which the rate declined with time as
292 the ripples grew in size.

293 The Zimbelman et al. (2009) and Jerolmack et al. (2006) ripple migration rates (Fig. 14) are notably
294 smaller than those we and Sharp (1963) have measured. The influence of sediment size or local
295 boundary layer properties is difficult to quantify, and it seems that freestream winds during our
296 observations were somewhat higher than those during the other work. One additional possibility, albeit
297 a seemingly unlikely one, is that some of these other measurements were 'aliased' by the long observing
298 interval. While the crest of a ripple was found some distance downwind of a marker placed at a crest
299 some hours before, it may be that it was not the same ripple. In other words, an apparent movement
300 X cm of a ripple pattern with wavelength λ may correspond to an actual movement of $X+n\lambda$, where
301 $n=0,1,2,3,\dots$. Non-zero and progressively larger values of n (and thus of the actual propagation speed)
302 are progressively less probable the shorter the measurement interval. Eliminating this possibility is one
303 advantage of the timelapse imaging approach.

304

305 4.2 Wind Statistics and Threshold

306 It is evident from the video record and Fig. 8 that ripple movement occurs only during a small fraction of
307 time. In part, this is due to winds simply being too weak to cause saltation and reptation, although
308 sometimes moisture (or, indeed, burial by snow !) may suppress movement even if winds are strong
309 enough.

310 Fig. 16 plots the wind gust data over some 7 years from the RAWS station (which may or may not -
311 depending on one's interpretation of Figs. 12 and 13 - underpredict winds at the site). It is seen that
312 while ~5 m/s winds are considered as a typical threshold for ripple movement (e.g. Zimbelman et al.,
313 2009), such gusts should occur most days at Great Sand Dunes, and indeed gusts of 20 m/s occur some
314 50% of the time. Further observations covering the summer season are hoped to elucidate the extent to
315 which the contrast between the small number of movements and the large number of days on which
316 winds were strong enough is due to, for example, moisture effects, or instead due to daily gusts being
317 too short-lived to be a good indicator of transport.

318 As is always the case in aeolian studies, high resolution (in time) local wind measurements would be
319 desired, but such aspirations always confront logistical limitations of battery energy, data volume etc.
320 One approach would be to record the mean-cube windspeed over a day, which would more faithfully
321 capture the drift potential than either the peak gust - which might be a very ephemeral windspeed - or
322 the average windspeed (in fact it was an attempt to record such a value that led to the coding error in
323 our PICAXE logger - future efforts will seek to achieve our goal).

324 4.3 Lessons in Technique

325 The ease of measurement and the 'watchability' of the movie were considerably improved over our
326 initial experiments in Egypt (Lorenz, 2011) by virtue of the mounting on a relatively tall and rigid post.
327 This improved the perspective of the ripple motion, and prevented scene judder due to camera motion.

328 An interesting feature in the scene is the moving shadow of the mounting post. This shadow projected
329 on to the variably-sloped scene acts as a useful scalebar. (Additionally, the fact that the shadow's path
330 across the scene varies during the 70-day observation due to the changing solar declination is of interest
331 in educational applications of these data.)

332 The possibility that the dominant sediment type may evolve during the observation suggests that some
333 means of evaluating this would be useful. One possibility would be a sand trap. Alternatively, some
334 sort of additional instrumentation such as a saltation detector that has a size-dependent response, or a
335 close-up camera able to resolve the size of grains, could be installed.

336

337 5. Conclusions

338 We believe this is the first long-duration (months) continuous field measurement of wind ripple
339 migration. The timelapse imagery approach demonstrated here generates a prodigious amount of data
340 with little effort. Extensive studies beyond the scope of the present paper could be performed, and the
341 data are being archived with the National Park Service to allow other investigators to exploit it. One
342 approach that efficiently condenses certain types of imagery is the 'waterfall chart', from which we have
343 extracted a set of ripple migration rates as a function of ripple size, with these variables having an
344 approximately reciprocal relationship. The instantaneous migration rates we have recorded exceed
345 previously published ripple movement rates, although we note that ripple movement in fact only
346 occurred on about 15% of the days we observed. Further investigation using the methods reported here
347 may help understand the influence of short-term wind variability and moisture on transport in field
348 conditions.

349

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- 386

387 Figure Captions

388 Figure 1. Aerial view of the ripple study site. Inset shows the location of the field site in relation to the
389 main Great Sand Dunes National Park and Preserve area. The insert map image uses Landsat image
390 acquired in 2000 with 14 m resolution, courtesy of NASA. The dune image is from the National
391 Agriculture Imagery Project, operated by the USDA.

392 Figure 2. Field photo, looking northeast along the western arm of the parabolic dune. Great Sand Dunes
393 are visible at the left. The apex of the parabolic dune is in the center of the image, and the two
394 fenceposts (one where Camera 2 was installed) are seen to its left, with the lone fencepost (to which
395 camera 1 and the anemometer were attached) is towards the right. The wire fence shows the position
396 of the road which has been overrun by the dune. Some ripples are visible near the fencepost.

397 Figure 3. Close-up field photo of a ripple crest. A US 1-cent coin ('penny') has a diameter of 19.05 mm.
398 At the crest, granules are ~2 mm across, fining to ~1 mm and below between crests. Beneath the
399 granule layer, fine sand of ~0.2 mm is present.

400 Figure 4. Camera 1 and anemometer installation. Datalogger and battery boxes are at the base of the
401 fencepost, which is 15 cm in diameter.

402 Figure 5. Montage of images from camera 1. (a) Two days after the beginning of the image sequence.
403 Note the large ripples and the shadow of the anemometer near the middle of the day (NB timestamp is
404 EST). (b) 11 days later, this morning image shows little change in the ripple pattern, here highlighted by
405 frost that has preferentially been retained on the ripple crests. (c) At the beginning of the Nov 21
406 movement event. (d) Two hours later, there has been noticeable movement: note also the increased
407 sinuosity of the ripples and the formation of a transverse small-wavelength ripple pattern at the bottom
408 center of the image. (e) 80 minutes after (d) some motion is detectable and the post shadow has moved
409 out of the field of view. Note also that a small streamer attached to the stake in the center of the image
410 shows the local winds are going uphill, i.e. to the right in the image. (f) Two days later, after further
411 motion, the ripples have become very discrete, with granule-free 'interdunes' and a much longer-
412 wavelength sinuosity. (g) In this image, the ripple pattern, now rather less sinuous and discrete, is
413 highlighted by snow between the crests. (h) Here, close to winter solstice, the shadow of the post is
414 much longer than in (a). (i) near the end of the sequence, the wavelength appears to have shortened.

415 Figure 6. Temperature history of the aluminium case of the anemometer datalogger. Note the low
416 temperatures encountered. Note also the period around day 43 where snow covered the scene and
417 pinned the temperature at the freezing point.

418 Figure 7. Peak wind gust in each one-hour period recorded by the anemometer shown in Fig. 4.

419 Figure 8. Ripple movement estimated by eye from Camera 1 imagery. The movements that are
420 discussed in detail in this paper are in the 19th day of the record.

421 Figure 9. Scene from camera 2, showing the fencepost and ripple pattern. The three dark lines are (top
422 to bottom) the image lines from which the three panels waterfall chart (Fig. 11) is constructed.

423 Figure 10. Unsharp mask / thresholded version of Fig. 9, used to construct the waterfall chart (Fig. 11).
424 The ripple pattern becomes sharply defined as a set of 'zebra stripes'.

425 Figure 11. Waterfall chart. Panels (a,b,c) correspond to the three image lines in Fig. 9, top to bottom.
426 Each panel shows contrast-enhanced line history from 10.03 to 17.33 hrs. Features A-K are features in
427 the chart from which speed measurements are extracted and reported in Fig. 14. Feature B is the
428 fencepost.

429 Figure 12. Windspeed history on 21 November, when the motions in figure 11 were recorded. The solid
430 line is the instantaneous windspeed (15 min intervals) at the Indian Springs station, while the dotted line
431 shows the peak hourly wind recorded by our anemometer at the site. The dotted and dot-dashed line
432 are the daily average and daily gust values from the RAWS station.

433 Figure 13. Scatterplot of the in-situ winds measured at the site (peak gust each day) against that
434 recorded at the RAWS station. Note the generally-higher winds at the site.

435 Figure 14. Ripple migration rates as a function of width as determined from the waterfall chart (Fig.
436 11). Grey ellipses marked J and Z correspond to measurements at other sites by Jerolmack et al. (2006)
437 and Zimbelman et al. (2009) respectively.

438 Figure 15. As for Fig. 14, but with the ripple width+spacing (i.e. 'wavelength') as the abscissa : the
439 scatter is much higher than for the width, suggesting that the width may be a tighter control on the
440 dynamics.

441 Figure 16. Probability distribution of peak daily gust exceeding a value V at the RAWS site, for which a
442 long record of some 7 years is available. Left panel shows linear axes, indicating gusts exceed ~ 20 m/s
443 half the time. Right panel has a logarithmic ordinate to highlight the occasional strongest gusts - gusts of
444 some 50 m/s are encountered every ~ 3 years. The dashed line indicates the high-speed gust data follow
445 an exponential fall-off of frequency.